1	Modeling the wake of the Marquesas archipelago
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19	Key Points:
20	• Characterizing the eddy activity in the island wake of the Marquesas archipelago using
21	high-resolution ocean model
22 23	• Revealing the importance of the fine-scale dynamics within the Island Mass Effect

24 Abstract

25 In this study, a high-resolution (~ 2.5 km) numerical model was set up to investigate the finescale activity within the region of the Marquesas archipelago. This has never been performed 26 before. The robustness of the model results is assessed by comparison with remote sensing and in 27 28 situ observations. Our results highlight regions of warm waters leeward of the different islands 29 with high eddy kinetic energy (EKE) on their sides. The analysis of energy conversion terms 30 reveals contributions to *EKE* variability by wind, baroclinic and barotropic instabilities. The use of a geometry-based eddy detection algorithm reveals the generation of cyclonic and anticyclonic 31 32 eddies in the wake of the largest islands, with both an inshore and offshore effect. Maximum 33 eddy activity occurs in austral winter following the seasonality of both wind stress and EKE 34 intensity. Most eddies have a radius between 20 and 30 km and are generally cyclonic rather than 35 anticyclonic. Significant vertical velocities are observed in the proximity of the islands, associated with topographically-induced flow separation. Eddy trapping inshore waters are 36 37 advected offshore in the wake of the islands. The overall influence of these fine-scale dynamics 38 could explain the strong biological enhancement of the archipelago.

39 1 Introduction

The Marquesas archipelago (144°W-137°W; 8°S-11°S) is located in the northern part of French 40 41 Polynesia, central South Pacific, where the South Equatorial Current (SEC) flows south-42 westward (Figure 1a). It is composed by a dozen of small volcanic islands with mountains up to 43 1224 m, rugged steep cliffs and no surrounding coral reefs (Savanier et al., 2006; Figure 1b). The main five islands are Nuku Hiva (339 km²), Ua Pou (105 km²) and Ua Huka (83 km²) in the 44 northern part of the archipelago, and Hiva Oa (320 km²) and Fatu Hiva (85 km²) in the southern 45 46 part. Despite their relatively small area coverage, a remarkable plume of chlorophyll-a (Chl, a proxy of phytoplankton biomass) can be observed leeward of the islands in the open ocean from 47 48 satellite derived ocean color [Signorini et al., 1999; Martinez and Maamaatuaiahutapu, 2004]. Such biological enhancement is referred to as an island mass effect (IME) [Doty, M. S., & Oguri, 49 50 1956]. While several physical processes can be involved in the present IME, such as coastal 51 upwelling, Ekman pumping, eddies or internal waves [Heywood et al., 1990, 1996; Barton, 52 2001; Sangrà et al., 2001; Palacios, 2002], the IME dominant mechanisms have not been 53 elucidated yet for the Marquesas archipelago.

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Figure 1: (a) Mean surface current from the satellite derived Ocean Surface Current Analysis – Real time (OSCAR) product (time averaged over Oct-1992 to Jun-2015, in m s⁻¹) over French Polynesia. The islands are represented in white, as well as the name of the five archipelagos. The black boxes show the parent and child grids implemented in the ROMS configuration. (b) Bottom topography from the 2-arc minute topography/bathymetry dataset ETOPO2 (in m) around the archipelago used in the model configuration as well as the names of the main islands.

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63 It is well known that oceanic currents encountering islands generate complex wakes. Several 64 studies in other regions have assessed these island wakes and their forcing mechanisms using a 65 remote-sensing [Wolanksi et al., 1996; DiGiacomo and Holt, 2001], in situ data [Hasegawa et al., 2004, 2008] or a modeling approach [Dietrich et al., 1996; Wolanksi et al., 1996; Dong et al., 66 2007; Jiménez et al., 2008; Hasegawa et al., 2009]. An important distinction has to be made 67 68 between shallow- and deep-water wakes. Shallow-water wakes occur when considering islands 69 in shallow shelves or estuaries where near-shore bottom drag acts as the primary source of 70 vorticity generation [Wolanksi et al., 1996; Alaee et al., 2004; Neill and Elliott, 2004]. Deep-71 water wakes occur leeward tall islands surrounding by a deep bathymetry - like in the Marguesas 72 archipelago - where bottom influence can be neglected and topographic and wind forcing are the 73 primary sources of vorticity generation. Topographic forcing refers to the detachment of the 74 frictional layer around the island as a result of differential bottom stress imparted to the flow by 75 the sloping sides of the island [Dong et al., 2007]. Wind forcing refers to eddy generation as a 76 consequence of Ekman pumping induced by wind shear in the island's wake [Jiménez et al., 77 2008].

78 Eddy formation and propagation in island wakes could play a key role in the IME of the 79 Marquesas archipelago. Indeed, a wind curl dipole has been reported in the archipelago 80 [Martinez et al., 2009] possibly forming oceanic cyclonic eddies and jets inducing an enrichment 81 of the upper layer leeward the islands by mechanisms such as the ones proposed by Hasegawa et 82 al. [2009] and Andrade et al. [2014]. Nevertheless, the previous studies on the Marquesas IME 83 using satellite altimetry did not mention any eddy features [Signorini et al., 1999; Legeckis et al., 84 2004; Martinez and Maamaatuaiahutapu, 2004]. This is probably due to the coarseness of the 85 Sea Level Anomaly (SLA) products used in these studies. However, some eddy activity does imprint the Sea Surface Temperature (SST). For example, on Nov. 17, 2015 (Figure 2a) a SST 86 87 dipole is observed leeward Nuku Hiva at (140.5W; 9°S), as well as a warm eddy further 88 southwest, while a cool eddy is present at (141°W; 10.1°S) on Aug. 21, 2012 (Figure 2b). In both 89 cases, it is important to note that the size of the structures is of the same order than the island 90 diameter (~30 km for Nuku Hiva). Such an eddy activity could possibly be at the origin of 91 nutrient uplift from deep rich waters to the euphotic zone allowing the development of 92 phytoplankton. Because the spatio-temporal resolution of satellite altimetry is too coarse to investigate the small scales of such aforementioned eddy activity around the Marquesas 93 94 archipelago, recourse to high-resolution modeling taking into account the topographic forcing 95 (i.e., the presence of the islands) has been done by the present study. Documenting the generation and characteristics of the oceanic eddies should be viewed as a first step toward an 96

97 explanation of the origin of the surface layer biological enrichment in this archipelago.



Figure 2: SST (°C) from Geostationary Operational Environmental Satellite (GOES)/ Polar
Operational Environmental Satellite (POES) in the Marquesas archipelago for (a) 2015-NOV-17
and (b) 2012-AUG-21.

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103 The present work is organized as follows: Section 2 presents the data and the methodology. A comparison of the oceanic conditions between the model, satellite and *in situ* observations is

performed in section 3. In this section, eddy properties generated leeward the islands are also

106 characterized. Finally, section 4 concludes this study.

107 2 Data and Methods

108 2.1 Numerical model

109 Our model is based on the Regional Ocean Modeling System (ROMS)-AGRIF (Adaptive Grid Refinement in Fortran) code provided by Debreu et al. [2012] and Penven et al. [2007]. The 110 ROMS model is a split-explicit, free surface and terrain-following vertical coordinate oceanic 111 112 model [Shchepetkin and McWilliams, 2003; 2005]. The AGRIF version is especially suitable to 113 study regional scale since it has the ability to manage an arbitrary number of fixed grids and 114 embedding levels. We configure the two-way embedding procedure. It means that the parent grid provides the boundary conditions for the child grid, and that the solution of the child grid is used 115 116 to improve the larger scale parent grid solution, allowing a smooth, continuous interfacing 117 between grid levels [Debreu and Blayo, 2008; Debreu et al., 2012]. Therefore, we defined two embedded grids: 1) the parent grid extends from 137°W to 144°W and 6°S to 12°S; 2) the child 118 119 grid extends from 138°W to 143°W and 7.3°S to 11.3°S as shown in Figure 1a. The grid 120 refinement rate is 3, implying a 1/15° (~7km) and 1/45° (~2.5 km) grid resolutions for the parent and child grids, respectively. This allows a sufficient sampling of the island topography and of 121 122 the fine-scale dynamics, as the first baroclinic Rossby radius of deformation is about 120 km in 123 this region according to Chelton et al. [1998]. The internal (external) time stepping is set to 3600 (600) sec for the parent grid and 1200 (200) sec for the child grid. Both grids have 32 vertical 124 125 levels and the vertical s-coordinate is stretched for boundary layer resolution. The topography is 126 derived from the 2' resolution ETOPO2 database provided by NOAA-NGDC [Smith and 127 Sandwell, 1997]. The bathymetry field has been filtered to keep the slope parameter < 0.25128 [Beckmann and Haidvogel, 1993]. The K-profile parameterization (KPP) vertical mixing scheme 129 from Large et al. [1994] is used to parameterize vertical mixing processes.

All the external forcings of the ROMS simulations are based on monthly climatologies. At the surface, the heat and fresh water fluxes are extracted on a monthly $\frac{1}{2}^{\circ}$ grid from the

- 132 Comprehensive Ocean-Atmosphere dataset (COADS) [da Silva et al., 1994]. The wind forcing is issued from the QuikSCAT monthly climatology calculated over 1999-2009, on a 1/4° grid 133 134 [Lungu and Callahan, 2006]. ROMS is connected to the lateral boundaries by an active, implicit 135 and upstream-biased radiation condition [Marchesiello et al., 2001]. The boundary conditions 136 and initial state are based on the objectively analysed World Ocean Atlas 2013 (WOA13) 137 monthly climatology on a 1/4° grid [Locarnini et al., 2013; Zweng et al., 2013]. The inflow boundary conditions are nudged toward temperature, salinity, and geostrophic velocity fields. 138 139 Following Kersalé et al., [2011], the nudging timescale for inflow and outflow are set to 1 day 140 and 1 yr for the tracer fields and 3 days and 1 yr for the momentum fields. The geostrophic
- velocity is referenced to the 1000 m depth. The explicit lateral viscosity is null all over the domain, except in the sponge layer. The width of the nudging border is 50 km and the maximum viscosity value for the sponge layer is set to $1000 \text{ m}^2 \text{ s}^{-1}$.
- 144 The high resolution ROMS simulation, hereafter referred to as W13Q, is run over 10 years with
- outputs averaged every 2 days. Integrated physical properties show that a statistical equilibrium
- 146 is reached by the model after the first year of simulation. To avoid any impact of the spin up on
- 147 the model output and to ensure the robustness of our findings, we decided to also entirely remove
- 148 the third year and to focus our investigation on the outputs from year 4 to 10.

149 2.2 Datasets

- 150 We compared our numerical results with several datasets. The Moderate Resolution Imaging
- 151 Spectroradiometer (MODIS) infrared SST measured from the Aqua satellite are available on a
- 152 monthly basis with a spatial resolution of 4 km. We use the time period from July 2002 to June
- 153 2015 to compute the monthly climatology.
- The 3-D monthly fields of temperature and salinity issued from the In Situ Analysis System (ISAS13) are based on Argo data [*Gaillard et al.*, 2016] and are available on a 0.5° grid from 2002 to 2012.
- 157 Monthly near surface current over 1992 to 2015 are obtained from the Ocean Surface Current 158 Analysis – Real time (OSCAR) with a 1/3° spatial resolution [*Bonjean and Lagerloef*, 2002].
- 159 Finally we used reanalyzed currents from the global Hybrid Coordinate Ocean model and the
- 160 Navy Coupled Ocean Data Assimilation (HYCOM+NCODA) [Cummings and Smedstad, 2013].
- 161 These velocity fields are available with a spatial resolution of $1/12.5^{\circ}$ and a daily resolution from 162 January 1st 2006 to December 31st 2012.
- 1632.3 Non-dimensional numbers
- 164 The ocean dynamics in island wake is generally characterized by the recourse to non-165 dimensional numbers issued from geophysical fluid dynamics. The wake is typically controlled 166 by the turbulent Reynolds number defined as:

$Re = U_0 D/\nu$

- 167 where U_0 is the unperturbed upstream velocity, D is the horizontal scale of the obstacle, and v is 168 the eddy viscosity [*Tomczak*, 1988]. To compute *Re* in the Marquesas Islands, we use an 169 horizontal eddy viscosity value of $v = 100 \text{ m}^2 \text{ s}^{-1}$ [*Heywood et al.*, 1990; *Jiménez et al.*, 2008]. 170 In the archipelago, only the biggest island, Nuku Hiva, exhibits a Reynolds number ($Re \approx 50$) 171 that exceeds the theoretical threshold of the Von Karman vortex street generation ($Re_{th} = 40$).
- To consider the spatial scale and energy involved in island wake, we also consider the Burger
- 173 number, representing the ratio between stratification and Earth's rotation, and defined as:

$$Bu = \left(\frac{R_d}{R}\right)^2$$

where *R* is the island radius and R_d the Rossby radius associated to the first baroclinic mode. Using the forcing data from WOA13, we found that the seasonality of R_d vary between120 and 130 km in the archipelago. When the island radius *R* is smaller than R_d ($Bu \ge 1$), a submesoscale wake is generated with eddies having their radius in the order of *R* [*Stegner*, 2014]. This is always the case in the archipelago where the radius of the largest island is about 20 km.

Finally, we calculate the geometric shallow-water parameter $\alpha = \frac{h}{R}$, with *h* the mean thermocline depth. We obtained an alpha ranging between 0.01 and 0.03. These values are typical of midocean isolated islands, such as Madeira [*Caldeira et al.*, 2002], Gran Canaria [*Sangrà et al.*, 2005] or Hawaii [*Calil et al.*, 2008], generating a deep-water wake.

183 2.4 Eddy detection and tracking

184 To detect eddies in our model experiment, the method based on the geometry of the flow field developed by Nencioli et al. [2010] is applied. This method is well established and used in 185 186 several studies with different resolution [Dong et al., 2012; Liu et al., 2012; Amores et al., 2013; 187 Mkhinini, N. et al., 2014]. It relies on the spatial definition of vortices defined as a region with a 188 rotary flow around its center. In other words, a vortex is defined as a region with velocity vectors 189 rotating around a center with a minimum speed. Its boundary is defined by the largest closed 190 contour line of the local stream function around the center. The flexibility of the algorithm depends on two parameters: a, the number of grid points where the current increases, and b, the 191 192 dimension of the area used to define the local minimum of velocity. These parameters set the 193 minimum size of the detectable vortices and allow the algorithm to work on different grid 194 resolution. Using the same protocol than Nencioli et al. [2010] and Liu et al. [2012], optimal 195 performances of the algorithm have been obtained with a = 2; b = 2 for our model 196 configuration and a = 3; b = 2 for the HYCOM reanalysis outputs.

197 Once eddies have been detected at each time step, we used the tracking method proposed by 198 Doglioli et al. [2007] and Nencioli et al. [2010]: the position and sign of each eddy center are 199 compared at successive time steps. An eddy track is identified when the eddy center at the next 200 time step is found within a given searching area of the previous time step. The searching radius is 201 the product of the averaged eddy center displacement speed and the time interval between two 202 model outputs (i.e., 2 days). To avoid that eddies move further than the searching area within 203 successive time steps (thus preventing track splitting), we added a constant of 20 cm s⁻¹ 204 corresponding to the eddy-center displacement speed and representing the mean velocity of the background current. Moreover, eddies with a shorter lifetime than 4 days are not considered to 205 ensure the only detection of consistent structures only. 206

The dataset obtained after applying the detection includes: eddy center locations, eddy contours, polarities and radii. Time evolution of each eddy detected is also recorded and gives us information on their generation and ending.

210 2.5 Eddy Kinetic Energy budgets

The *EKE* is defined as the kinetic energy due to transient dynamic. Neglecting the vertical velocity contributions: $EKE = \frac{1}{2}(u'^2 + v'^2)$, where $u' = u - \bar{u}$ and $v' = v - \bar{v}$, with u and vthe zonal and meridian components of the velocity at 10 m, \bar{u} and \bar{v} their temporal means from year 4 to 10. To consistently compare our numerical model results with the monthly satellite 215 derived observations (OSCAR), we subsampled the W13Q current components at 10 m (u_m and 216 v_m) by extracting one grid point every 5 grid points. Then, we used monthly averages to 217 compute the monthly $EKE_m = \frac{1}{2}(u_m'^2 + v_m'^2)$.

- 218 To quantify the relative importance of instability and eddy-mean interaction mechanisms on the
- 219 EKE generation, we consider the different terms of energy budgets relative to the EKE
- 220 generation [Auad et al., 1991; Marchesiello et al., 2003; Liang et al., 2012]. The wind work, the
- barotropic and the baroclinic energy conversion [Dong et al., 2007; Halo et al., 2014; Hristova et
- *al.*, 2014; *Sun et al.*, 2016] are defined, respectively, as:

$$FeKe = -\frac{1}{\rho_0} \left(\overline{u'\tau_x'} + \overline{v'\tau_y'} \right)$$
$$KmKe = -\left(\overline{u'u'} \frac{\partial \overline{u}}{\partial x} + \overline{u'v'} \frac{\partial \overline{u}}{\partial y} + \overline{u'w'} \frac{\partial \overline{u}}{\partial z} + \overline{v'u'} \frac{\partial \overline{v}}{\partial x} + \overline{v'v'} \frac{\partial \overline{v}}{\partial y} + \overline{v'w'} \frac{\partial \overline{v}}{\partial z} \right)$$
$$PeKe = -\frac{g}{\rho_0} \overline{\rho'w'}$$

where τ_x and τ_y are the meridian and zonal component of the wind stress. The prime indicates 223 the anomaly from the annual mean. g is the gravity, ρ is the density and $\rho_0 = 1030$ kg m⁻³. The 224 overbar denotes the temporal mean and the prime represents the anomaly deviation from the 225 226 mean. The FeKe is the EKE generation due to transient wind. When positive, this term induces an energy input while, when negative, it implies a damping effect on the sea surface [Xu and 227 228 Scott, 2008]. The KmKe is the energy conversion between mean currents and EKE. When its 229 volume-integration is positive, it implies barotropic energy conversion. The *PeKe* is the energy conversion between available potential energy and EKE and indicates baroclinic instability when 230 231 its volume-integration is positive [Harrison and Robinson, 1978]. Both KmKe and PeKe are 232 vertically integrated over the surface layer (top 100 m) [Hristova et al., 2014]. The minimum 233 depth in the model is set to 75 m. When the bathymetry is shallower than 100 m, conversion 234 terms are integrated over the actual depth.

235

3 Results and Discussion

237 3.1 Thermohaline structures and seasonal variability

238 The Marquesas archipelago is characterized by a north-east/south-west SST gradient as shown 239 by MODIS remote sensing data and W13Q (Figure 3a and 3b). Colder water in the north-east (≤ 240 27.5°C) represents the imprint of the equatorial upwelling [Wyrtki, 1981] while warmer SST in 241 the west ($\geq 28^{\circ}$ C) is the south-eastward imprint of the western Pacific warm pool [Yan et al., 242 1992]. The high-resolution of MODIS reveals warm regions behind the islands likely due to 243 areas zones of weak currents leeward the islands. These structures have also been observed 244 leeward several islands elsewhere [Caldeira and Marchesiello, 2002; Caldeira et al., 2002]. The maximum of the mean satellite SST over the archipelago is up to 28.6°C in April during the 245 246 austral summer (Figure 3c, black line). In austral winter (September), SST in the Marquesas 247 region decreases down to 27°C. In this tropical region, the seasonal amplitude is around 2°C as confirmed by in-situ SST from ISAS13 (red line). All the aforementioned features appear in the 248 249 same locations and with comparable amplitudes in our numerical experiment, although SST is 250 slightly colder than satellite observations (0.05°C to 0.1°C; Figure 3b vs. 3a; green line on Figure





Figure 3: Mean SST over the child grid area from (a) MODIS (time averaged from Jul-2002 to Jun-2015) and (b) W13Q at 10m (time averaged over year 4 to 10). (c) Monthly climatology averaged over the child grid area from MODIS (black), ISAS13 (red) and W13Q (green). The grey shaded area represents the spatial standard deviation from MODIS monthly climatology. Units are in °C.

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260 Considering the haline structure, in situ data from ISAS13 show a SSS maximum in austral 261 summer and a minimum in austral winter (35.7 and 35.5 psu +/- 0.01 psu), that are satisfactory reproduced by the model (Figure 4a). The TS diagram issued from ISAS13 and W13O illustrates 262 263 the water masses within the archipelago (Figure 4b). The South Pacific Tropical Water (SPTW) is characterized by a salinity maximum in subsurface (> 35.6 psu) [O'Connor et al., 2005]. The 264 265 Eastern South Pacific Central Water (ESPCW) and the Antarctic Intermediate Water (AAIW) are respectively defined by temperature ranges between 8 °C and 24 °C and between 2 °C and 10 °C. 266 Their salinities are ranging between 34.4 psu and 36.4 psu and between 33.8 psu and 34.5 psu, 267 respectively [Emery, 2001]. Only the SPTW is fresher in W13Q than with in situ measurements. 268 269 The other water masses are quite well reproduced.



Figure 4: (a) Time series of the monthly averaged salinity (in psu) at 10 m issued from ISAS13 (red) and W13Q (green). Red shaded area corresponds to the spatial standard deviation from ISAS13. (b) TS diagram from ISAS13 (red) and W13Q (green) monthly climatology and their standard deviations (red and green shades respectively).



276 3.2 Current structures

As shown in Figure 1a, the SEC is the main surface current flowing around the Marguesas 277 278 archipelago. Currents issued from W13Q present the same order of magnitude than the OSCAR 279 ones, although with a more westward direction in the north (Figure 5a). The spatial distribution 280 of the mean current in the archipelago can be depicted in two regions: the southern part where the horizontal current is weak (8 cm s^{-1}) , and the northern part where the SEC is stronger and 281 reaches values of 20 cm s⁻¹ as observed by *Martinez et al.* [2009]. The deviation of the SEC 282 creates stronger currents on both sides of the islands and regions of weak current just behind the 283 islands (Figure 5b) [Chang et al., 2013; Karnauskas et al., 2017]. These shadow zones of weak 284 285 currents are collocated with the warm areas as reported previously in Figure 3b.





Figure 5: (a) W13Q current at 10 m time averaged over year 4 to 10 and over the child domain. For clarity, only 1 vector out of 17 is represented. (b) Zoom of (a) over the Marquesas northern islands, as defined by the black box on a). For clarity, only 1 vector out of 3 is represented. Units

are in m s⁻¹.

293 3.3 *EKE* generation and seasonal variability

294 In order to investigate the dynamical properties within the archipelago and more precisely the eddy field, we first focus on the EKE characterisation. The EKE_m obtained from OSCAR 295 monthly currents show homogeneously low values around the archipelago $(25 \text{ cm}^2 \text{ s}^{-2})$ in Figure 296 6a. In the northern and the south-western part of the archipelago, two areas are barely more 297 298 active, the former corresponding to the southward imprint of the turbulent equatorial area [Oiu 299 and Chen, 2004]. To compare the eddy activity from W13Q with the remote sensed one, we subsampled the model outputs to the same spatial and temporal resolution before computing 300 EKE_m (Figure 6b). High activity occurs in the northern and south-western part of the 301 302 archipelago. Monthly W13Q also presents a high eddy activity west of Nuku Hiva. To fully take 303 advantage of the high-resolution model, the annual mean of *EKE* is now directly computed from the W13Q two-day outputs (Figure 6c). Thanks to the higher spatial and temporal resolution, the 304 *EKE* at 10 m reaches up to 90 cm² s⁻² (vs. 55 cm² s⁻² in Figure 6b). Island wakes characterized by 305 a highly variable EKE pattern are highlighted westward the northern islands of Nuku Hiva, Ua 306 307 Pou and Ua Huka. We could have also expected high activity leeward Hiva Oa (in the south), 308 since its diameter is approximately the same as Nuku Hiva. However, the SEC being weaker 309 windward this island, it does not produce a strong enough EKE. Consistently, EKE patterns are 310 more pronounced in the northern part of the archipelago where the mean current is stronger as seen on Figure 5a. The EKE standard deviations are given in Figure 6e and 6f. They reveal high 311 312 activity where the mean EKE is already high and in good agreement with OSCAR EKE standard 313 deviations (Figure 6d).



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Figure 6: Annual mean of EKE_m (in cm² s⁻²) obtained from monthly averaged of (a) OSCAR surface currents and (b) W13Q currents at 10m. (c) is the annual mean of *EKE* derived from the W13Q 2-day outputs. Their standard deviations are represented in (d), (e) and (f), respectively.

319 In order to determinate the *EKE* origin, we now investigate the energy budget. For all the islands, the Burger number Bu > 1. Therefore, a significant amount of kinetic energy in the *EKE* origin is 320 321 expected. This trend is observed in the barotropic energy conversion (Figure 7a). The *KmKe* is 322 particularly intense leeward Nuku Hiva where the EKE pattern is the strongest (Figure 6c). It is 323 mainly due to the fact that Nuku Hiva is the largest island and that it lies where the currents are 324 the strongest. This implies a Re = 58 and, consequently, a possible Von Karman wake 325 generation. KmKe and PeKe spatial patterns are close and are driven by the topography of the 326 archipelago (Figure 7a, b and Figure 1b). The distribution of the baroclinic energy conversion is 327 negative in average, which indicates an eddy dissipation toward the eddy potential energy [Kang 328 and Curchitser, 2015]. The averaged energy input from the wind (FeKe) presents smooth 329 patterns because of the relatively low resolution of the wind stress used in the model (Figure 7c). 330 Values are weaker than the two previous parameters.

331 The *EKE* seasonality is superimposed with the different energy budget terms and their sum in 332 Figure 7d. EKE time series show a maximum activity in February, June and August, when the 333 *KmKe* is also high. The latter remains positive all year round, indicating a constant flux towards 334 EKE. PeKe, in a smaller proportion, also contributes to the high EKE activity in February. PeKe 335 is responsible for the minimum value of EKE in April. The EKE variation on the second half of the year is explained by the *FeKe* variation, with a decreasing activity from August to December. 336 337 The sum of these energy flux terms is relatively well correlated with the EKE variability. When 338 this sum is positive (negative), an increase (decrease) of EKE is observed except during June 339 where the energy budget seems too low to explain the EKE activity. According to Chen et al. 340 [2016], the energy transfer through eddy-mean flow interactions could be imbalanced and 341 compensated by advection. Indeed, currents are relatively intense in the archipelago during this 342 time period. This could be at the origin of the decay observed in these time series. Although the 343 spatial averaging reveals low values of *PeKe* and *FeKe* compared to some local maxima (Figure 344 7a and b), these energy inputs seem to play an important role in the EKE seasonal variation, 345 particularly at the beginning and the end of the year. 346



Figure 7: Spatial distribution of the annual mean of the (a) barotropic (*KmKe*), (b) baroclinic (*PeKe*) energy conversion both integrated over the surface to 100 m and (c) the *EKE* generation due to transient wind, from year 4 to 10 of W13Q (units are in cm³ s⁻³). (d) Time series of wind work, baroclinic and barotropic energy (left axis), in yellow, blue and red, respectively. The sum of these three terms is indicated in black. The *EKE* (in cm² s⁻², right axis) spatially averaged over the child grid and at 10 m is represented in green.

354 3.4 Eddy activity

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Detected eddies within the Marquesas archipelago are illustrated with green contours on relative 356 vorticity maps from June 22 to July 22 during year 9 of W13Q (Figure 8). Positive (negative) 357 358 values reveal anticyclonic (cyclonic) circulation. The total number of detected eddies over the 7 years of integration is 1260, with a dominance of cyclonic (714) vs. anticyclonic eddies (546). 359 360 This cyclonic eddy dominance (i.e., 56.7% of the total eddy number) is well illustrated (Figure 361 8). On the other hand, positive patterns are unstable and tend to broaden and fragment inducing a weaker proportion of anticyclonic eddies. These observations are consistent with previous 362 363 studies and are due to centrifugal instability occurring to anticyclonic eddies [Dong et al., 2007; 364 Hasegawa et al., 2009; Stegner, 2014]. These snapshots reveal an important eddy activity in the archipelago which has never been investigated before. In particular, Figure 8c shows the 365 366 generation of cyclonic and anticyclonic eddies leeward Nuku Hiva, the northern island, and likely producing SST small-scale patterns as observed in Figure 2. 367



Figure 8: Snapshots of relative vorticity at 10 m (units are in s⁻¹) from W13Q during Year 9 every 6 days (a-f). The detected eddies are represented by *green* contours (Nencioli et al. [2010] based algorithm). The animation of the vorticity from June to July is available in the auxiliary material.

374 To investigate where eddies are generated, the first eddy position determined by the detection 375 algorithm is recorded and analysed. Most of eddies are generated leeward the archipelago 376 highlighting the role of the topography (Figure 9a). Part of these eddies are generated just behind the islands. Although the mean Reynolds number is larger than 40 only for Nuku Hiva, eddies 377 378 are also generated behind the other islands. Indeed, when surface currents accelerate on the 379 flanks of the other islands (Figure 5), they induce a larger Re and create a decreasing pressure in 380 the lee side associated with two opposite recirculation cells behind the islands. Anticyclonic 381 eddies are generated on the equatorward side and cyclonic eddies on the poleward side. These 382 eddies are emphasized behind Nuku Hiva and Ua Pou in the northern part of the archipelago 383 where the SEC is the strongest, and Hiva Oa in the south which is the second biggest island of 384 the Marquesas archipelago. Conversely, there is no eddy generation behind Ua Huka (the 385 northeastern most island) due to its small size inducing a too low Re (= 34) while there is a 386 signature of the island wake on EKE (Figure 6c), mean currents (Figure 5), energy conversion 387 rates (Figure 7a) and vorticity (Figure 8).

Eddies are also generated farther from the coasts, in the open ocean, where the flow presents complex structures as seen on Figure 8. They are generated almost exclusively south-westward of the islands. A small number is also generated elsewhere, likely due to an eastward perturbation of the SEC during austral winter and referred as the Marquesas Counter Current [*Martinez et al.*, 2009]. Most of the eddies in the archipelago are generated during austral winter (July and August) when the wind stress and the *EKE* are maximum (Figure 9b and Figure 7c). Contrarily, eddies are barely detected in March and April when the wind stress and *EKE* are minimum. A strong correlation is found between eddy generation and the wind stress, with r = 0.87 ($p = 2.10^{-4}$). A high eddy generation occurs when the wind is strong (Figure 9b). Indeed, not only the wind is responsible for a part of the *EKE* generation, but it also strengthens the surface currents, providing sufficient velocity to generate eddies behind the islands.



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Figure 9: (a) Location of eddy generation identified from W13Q over year 4 to 10. Anticyclonic
and cyclonic eddies are represented by red and blue circles, respectively. (b) Monthly mean
number of generated eddies from W13Q from year 4 to 10 (green line) and wind stress issued
from QuickSCAT in N.m⁻² (orange line).

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405 Investigating the maximum size of the detected eddies during their lifetime in W13Q shows that 406 their distribution is similar to a Gaussian without an inferior tail due to resolution limitations 407 (Figure 10a, green bars). Most of the eddy radii are lower than 25 km and only a few are larger 408 than 45 km. Indeed, size of eddies generated leeward the islands are approximately the same than 409 the island dimensions, which can be explained by Bu larger than 1. The HYCOM eddy size distribution from January 2006 to December 2012 has been investigated within the child grid 410 411 (Figure 10a, black bars). In opposition with the W13Q results, a significant number of eddies 412 larger than 20 km is present while smaller eddies are absent in HYCOM. The coarser spatial 413 resolution of HYCOM prevents the resolution of such small structures. HYCOM larger eddies 414 are suspected to be induced by the inter-annual forcing which is not represented in the present 415 climatological W13Q configuration. In total, almost three times more eddies, are detected in our 416 configuration than in HYCOM resulting from the higher resolution of our model.

417 Considering now the cyclonic *vs.* anticyclonic patterns, W13Q cyclonic eddies are dominant for 418 all radius classes, except for the smallest ones (~10km) (Figure 10b). Nevertheless, as previously



419 mentioned, the generated anticyclonic structures are more unstable and form small structures420 with short lifetime (shorter than a week).

Figure 10: (a) Distribution of the detected eddy radius from year 4 to 10 in W13Q (green) and from 2006 to 2012 in HYCOM (black). (b) Distribution of W13Q cyclonic (blue) vs. anticyclonic (red) eddies.

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Focusing on these small eddies of interest, we further investigate trajectories of eddies with a 426 427 maximum radius between 20 and 30 km, which represents about 50% of the eddies (Figure 10a). 428 As presented in Figure 10b, anticyclonic eddies are less numerous than cyclonic ones for this 429 radius range (Figure 11a vs. 11b, respectively). Furthermore, the anticyclonic eddy trajectories 430 are shorter than the cyclonic ones (Figure 11b) because of the centrifugal instability that weakens 431 anticyclonic eddies [Dong et al., 2007]. While Calil et al. [2008] showed that cyclonic 432 (anticyclonic) eddies tend to propagate poleward (equatorward) in the Hawaiian archipelago, 433 trajectories of eddies generated leeward the Marquesas Islands are mostly zonal and do not 434 present such a meridian trend. Indeed, while 61% (54.8%) of the cyclonic (anticyclonic) eddies 435 propagate equatorward (southward), their meridian mean latitude change is only of 0.0163° 436 (0.0182°). This might be explained by the lower latitude of the Marquesas archipelago inducing a 437 weaker beta effect than in Hawaii.



Figure 11: Eddy generation sites (black circles) and their trajectories for (a) anticyclonic and (b)
cyclonic eddies detected in W13Q and for a maximum radius over 20-30 km encountered over
their lifetime. The size of the radius along the trajectory is given by the color bar (in km).

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3.5 An insight to a possible eddy induced biological activity

444 Hasegawa et al. [2009] hypothesized that a combination of three processes can enrich the 445 surface layer in island wakes: the propagation of coastal rich waters formed in the lee of the islands by eddy shedding, the vertical advection of the Deep Chlorophyll Maximum (DCM) 446 447 toward the surface in the island wake, and the uplift of nutrient rich deep waters by the strongest cyclonic eddies. We show that, when the SEC encounters the Marquesas Islands, it generates 448 449 eddies in the lee of the archipelago (Figure 9a). These eddies can trap and propagate inshore 450 waters in the island wake. Thus, if these inshore waters have been previously enriched in 451 proximity of the islands, the biological activity can be enhanced. An increase of the sea surface 452 concentration of chlorophyll in the lee of an island could also be induced by the vertical 453 advection of the DCM. During the "Pakaihi i te Moana" expedition, a correlation between the depth of the DCM and the mixed layer depth (MLD) has been reported in the archipelago 454 [Martinez et al., 2016]. On average, the DCM is deeper (shallower) in the northern (southern) 455 456 region of the archipelago consistently with the MLD (60 m vs. 30 m). Our results show strong 457 vertical velocities at the MLD associated with the eddies of the island wake (Figure 12). This 458 high vertical activity might induce the vertical advection of the DCM toward the surface. The 459 evaluation of the impact of the fine scale dynamics on the nutricline is beyond the scope of the present study and will need further research. Indeed, to assess the uplift of nutrient-rich deep 460 461 waters by cyclonic eddies and the associated biological activity, a coupled physical-462 biogeochemical model is required. 463



Figure 12: Snapshots of vertical velocities (in m day⁻¹) as in Figure 8, but at the mixed layer depth.

467

468 **4 Conclusions and perspectives**

The island wake in the Marquesas archipelago has been characterized in the present study, providing new insights into the spatial and seasonal variability of the fine-scale dynamics of this oceanic region. We based our analysis on the ocean circulation model ROMS-AGRIF with a 1/45° child grid resolution and forced by climatological mean fields. The high spatial and temporal resolution allows characterizing warm SST patterns just behind the Marquesas Islands. These patterns likely due to shadow zones of weak currents have been consistently observed leeward several islands [*Barton*, 2001; *Caldeira et al.*, 2002; *Chavanne et al.*, 2002].

476 Accordingly to a rough approximation of the local Reynolds number, only Nuku Hiva, the 477 biggest island of the archipelago, would be expected to generate a wake of propagating eddies.

478 However, the mean *EKE* and the use of an automated eddy detection algorithm reveals that eddy

479 generation also occurs leeward the other islands, inshore and offshore. The energy budget 480 analysis reveals that eddies are generated by a combination of wind power input, barotropic and

481 baroclinic energy conversion. These generated eddies are cyclonic dominant due to the small size

482 of the islands in comparison to the Rossby radius. Most of the eddies have their maximum radius
483 between 20 and 30 km.

484 Our physical model results provide insight to explain the Marquesas biological enhancement. We

485 suggest that the combination of at least two of the three processes proposed by *Hasegawa et al.*

486 [2009] enriches surface layers via eddy activity: the propagation of rich waters formed in the lee

487 of the islands by eddy shedding and the vertical advection of the DCM toward the surface in the

island wake. To further investigate the third process, I.E. the possibility of an uplift of nutrient
rich waters by the cyclonic eddies, new in situ data and the use of a coupled physicalbiogeochemical numerical model are required.

491

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496 The data used in this work are available from Moderate Resolution Imaging Spectroradiometer 497 (http://oceandata.sci.gsfc.nasa.gov), MODIS In Situ Analysis System ISAS13 498 (http://www.seanoe.org/data/00348/45945/), Ocean Surface Current Analysis - Real time 499 OSCAR (http://www.oscar.noaa.gov/) and Hybrid Coordinate Ocean Model HYCOM 500 (http://hycom.org/data/glbu0pt08/expt-19pt1). Please contact the corresponding author to obtain 501 the data of the numerical model used in this paper.

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Figure 1: (a) Mean surface current from the satellite derived Ocean Surface Current Analysis – Real time (OSCAR) product (time averaged over Oct-1992 to Jun-2015, in m s⁻¹) over French Polynesia. The islands are represented in white, as well as the name of the five archipelagos. The black boxes show the parent and child grids implemented in the ROMS configuration. (b) Bottom topography from the 2-arc minute topography/bathymetry dataset ETOPO2 (in m) around the archipelago used in the model configuration as well as the names of the main islands.

Figure 2: SST (°C) from Geostationary Operational Environmental Satellite (GOES)/ Polar Operational Environmental Satellite (POES) in the Marquesas archipelago for (a) 2015-NOV-17 and (b) 2012-AUG-21.

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Figure 3: Mean SST over the child grid area from (a) MODIS (time averaged from Jul-2002 to Jun-2015) and (b) W13Q at 10m (time averaged over year 4 to 10). (c) Monthly climatology averaged over the child grid area from MODIS (black), ISAS13 (red) and W13Q (green). The grey shaded area represents the spatial standard deviation from MODIS monthly climatology. Units are in °C.

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Figure 4: (a) Time series of the monthly averaged salinity (in psu) at 10 m issued from ISAS13
(red) and W13Q (green). Red shaded area corresponds to the spatial standard deviation from
ISAS13. (b) TS diagram from ISAS13 (red) and W13Q (green) monthly climatology and their
standard deviations (red and green shades respectively).

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Figure 5: (a) W13Q current at 10 m time averaged over year 4 to 10 and over the child domain.
For clarity, only 1 vector each 17 is represented. (b) Zoom of (a) over the Marquesas northern
islands, as defined by the black box on a). For clarity, only 1 vector each 3 is represented. Units
are in m s⁻¹.

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Figure 6: Annual mean of EKE_m (in cm² s⁻²) obtained from monthly averaged of (a) OSCAR surface currents and (b) W13Q currents at 10m. (c) is the annual mean of *EKE* derived from the W13Q 2-day outputs. Their standard deviations are represented in (d), (e) and (f), respectively.

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Figure 7: Spatial distribution of the annual mean of the (a) barotropic (*KmKe*), (b) baroclinic (*PeKe*) energy conversion both integrated over the surface to 100 m and (c) the *EKE* generation due to transient wind, from year 4 to 10 of W13Q (units are in cm³ s⁻³). (d) Time series of wind

- work herealinia and heretronia energy (left axis) in valley, hive and red regreatively. The sum
- 737 work, baroclinic and barotropic energy (left axis), in yellow, blue and red, respectively. The sum

of these three terms is indicated in black. The *EKE* (in cm² s⁻², right axis) spatially averaged over the child grid and at 10 m is represented in green.

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Figure 8: Snapshots of relative vorticity at 10 m (units are in s⁻¹) from W13Q during Year 9 every 6 days (a-f). The detected eddies are represented by *green* contours (Nencioli et al. [2010]

- based algorithm). The animation of the vorticity from June to July is available in the auxiliary material.
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Figure 9: (a) Location of eddy generation identified from W13Q over year 4 to 10. Anticyclonic
and cyclonic eddies are represented by red and blue circles, respectively. (b) Monthly mean
number of generated eddies from W13Q from year 4 to 10 (green line) and wind stress issued
from QuickSCAT in N.m⁻² (orange line).

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Figure 10: (a) Distribution of the detected eddy radius from year 4 to 10 in W13Q (green) and from 2006 to 2012 in HYCOM (black). (b) Distribution of W13Q cyclonic (blue) vs. anticyclonic (red) eddies.

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Figure 11: Eddy generation sites (black circles) and their trajectories for (a) anticyclonic and (b) cyclonic eddies detected in W13Q and for a maximum radius over 20-30 km encountered over

their lifetime. The size of the radius along the trajectory is given by the color bar (in km).

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Figure 12: Snapshots of vertical velocities (in m day⁻¹) as in Figure 8, but at the mixed layer depth.